

INTERNATIONAL STANDARD

ISO
27852

Second edition
2016-07-01

Space systems — Estimation of orbit lifetime

Systèmes spatiaux — Estimation de la durée de vie en orbite

iTeh Standards
(<https://standards.iteh.ai>)
Document Preview

[ISO 27852:2016](#)

<https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016>



Reference number
ISO 27852:2016(E)

© ISO 2016

iTeh Standards
(<https://standards.iteh.ai>)
Document Preview

[ISO 27852:2016](#)

<https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016>



COPYRIGHT PROTECTED DOCUMENT

© ISO 2016, Published in Switzerland

All rights reserved. Unless otherwise specified, no part of this publication may be reproduced or utilized otherwise in any form or by any means, electronic or mechanical, including photocopying, or posting on the internet or an intranet, without prior written permission. Permission can be requested from either ISO at the address below or ISO's member body in the country of the requester.

ISO copyright office
Ch. de Blandonnet 8 • CP 401
CH-1214 Vernier, Geneva, Switzerland
Tel. +41 22 749 01 11
Fax +41 22 749 09 47
copyright@iso.org
www.iso.org

Contents

Page

Foreword	iv
Introduction	v
1 Scope	1
2 Normative References	1
3 Terms, definitions, symbols and abbreviated terms	1
3.1 Terms and definitions	1
3.2 Symbols	3
3.3 Abbreviated terms	4
4 Orbit lifetime estimation	4
4.1 General requirements	4
4.2 Definition of orbit lifetime estimation process	4
5 Orbit lifetime estimation methods and applicability	5
5.1 General	5
5.2 Method 1: High-precision numerical integration	6
5.3 Method 2: Rapid semi-analytical orbit propagation	7
5.4 Method 3: Numerical table look-up, analysis and fit formula evaluations	7
5.5 Orbit lifetime sensitivity to sun-synchronous	7
5.6 Orbit lifetime statistical approach for high-eccentricity orbits (e.g. GTO)	7
6 Drag modelling	13
6.1 General	13
6.2 Atmospheric density modelling	13
6.3 Long-duration solar flux and geomagnetic indices prediction	14
6.4 Approach 1: Monte Carlo random draw of solar flux and geomagnetic indices	15
6.5 Method 3: Equivalent constant solar flux and geomagnetic indices	19
6.6 Atmospheric density implications of thermospheric global cooling	23
7 Estimating ballistic coefficient (C_{DA}/m)	23
7.1 General	23
7.2 Estimating aerodynamic force and SRP coefficients	24
7.2.1 Aerodynamic and solar radiation pressure coefficient estimation via a “panel model”	24
7.3 Estimating cross-sectional area with tumbling and stabilization modes	27
7.4 Estimating mass	28
Annex A (informative) Space population distribution	29
Annex B (informative) 25-year lifetime predictions using random draw approach	32
Annex C (informative) Solar radiation pressure and 3rd-body perturbations	37
Annex D (informative) Sample code for drag coefficient estimation via panel model	39
Bibliography	41

Foreword

ISO (the International Organization for Standardization) is a worldwide federation of national standards bodies (ISO member bodies). The work of preparing International Standards is normally carried out through ISO technical committees. Each member body interested in a subject for which a technical committee has been established has the right to be represented on that committee. International organizations, governmental and non-governmental, in liaison with ISO, also take part in the work. ISO collaborates closely with the International Electrotechnical Commission (IEC) on all matters of electrotechnical standardization.

The procedures used to develop this document and those intended for its further maintenance are described in the ISO/IEC Directives, Part 1. In particular the different approval criteria needed for the different types of ISO documents should be noted. This document was drafted in accordance with the editorial rules of the ISO/IEC Directives, Part 2 (see www.iso.org/directives).

Attention is drawn to the possibility that some of the elements of this document may be the subject of patent rights. ISO shall not be held responsible for identifying any or all such patent rights. Details of any patent rights identified during the development of the document will be in the Introduction and/or on the ISO list of patent declarations received (see www.iso.org/patents).

Any trade name used in this document is information given for the convenience of users and does not constitute an endorsement.

For an explanation on the meaning of ISO specific terms and expressions related to conformity assessment, as well as information about ISO's adherence to the WTO principles in the Technical Barriers to Trade (TBT) see the following URL: [Foreword - Supplementary information](#)

The committee responsible for this document is ISO/TC 20, *Aircraft and space vehicles*, Subcommittee SC 14, *Space systems and operations*.

This second edition cancels and replaces the first edition (ISO 27852:2011), which has been technically revised.

[ISO 27852:2016](#)

<https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016>

Introduction

This International Standard is a supporting document to ISO 24113 and the GEO and LEO disposal standards that are derived from ISO 24113. The purpose of this International Standard is to provide a common consensus approach to determining orbit lifetime, one that is sufficiently precise and easily implemented for the purpose of demonstrating compliance with ISO 24113. This project offers standardized guidance and analysis methods to estimate orbital lifetime for all LEO-crossing orbit classes.

iTeh Standards
(<https://standards.iteh.ai>)
Document Preview

[ISO 27852:2016](#)

<https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016>

Space systems — Estimation of orbit lifetime

1 Scope

This International Standard describes a process for the estimation of orbit lifetime for spacecraft, launch vehicles, upper stages and associated debris in LEO-crossing orbits.

This International Standard also clarifies the following:

- a) modelling approaches and resources for solar and geomagnetic activity modelling;
- b) resources for atmosphere model selection;
- c) approaches for spacecraft ballistic coefficient estimation.

2 Normative References

The following documents, in whole or in part, are normatively referenced in this document and are indispensable for its application. For dated references, only the edition cited applies. For undated references, the latest edition of the referenced document (including any amendments) applies.

ISO 24113, *Space systems — Space debris mitigation requirements*

3 Terms, definitions, symbols and abbreviated terms

3.1 Terms and definitions

For the purposes of this document, the terms and definitions given in ISO 24113 and the following apply.

<https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016>

3.1.1 orbit lifetime

elapsed time between the orbiting spacecraft's initial or reference position and orbit demise/reentry

Note 1 to entry: An example of the orbiting spacecraft's reference position is the post-mission orbit.

Note 2 to entry: The orbit's decay is typically represented by the reduction in perigee and apogee altitudes (or radii) as shown in [Figure 1](#).

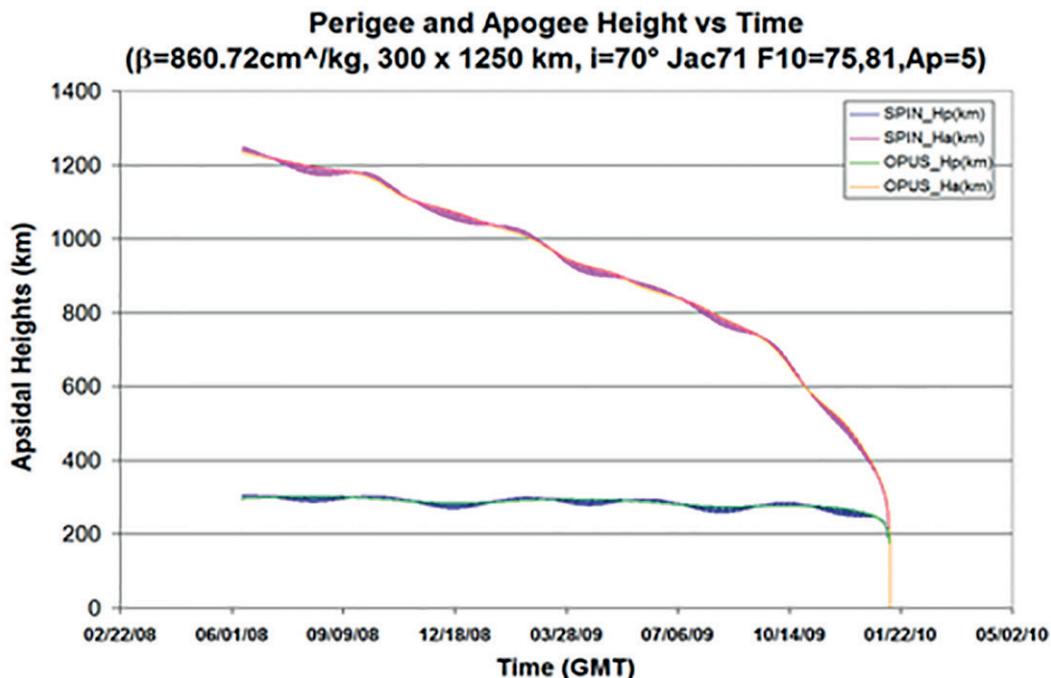


Figure 1 — Sample of orbit lifetime decay profile

3.1.2

earth equatorial radius

equatorial radius of the Earth

Note 1 to entry: The equatorial radius of the Earth is taken as 6 378,137 km and this radius is used as the reference for the Earth's surface from which the orbit regions are defined.

3.1.3

[ISO 27852:2016](#)

high area-to-mass

<https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016>

HAMR

space objects are considered to be high area-to-mass (or HAMR) objects if the ratio of area to mass exceeds 0,1 m²/kg

3.1.4

LEO-crossing orbit

low-earth orbit, defined as an orbit with perigee altitude of 2 000 km or less

Note 1 to entry: As can be seen in [Figure A.1](#), orbits having this definition encompass the majority of the high spatial density spike of spacecraft and space debris.

3.1.5

long-duration orbit lifetime prediction

orbit lifetime prediction spanning two solar cycles or more (e.g. 25-year orbit lifetime)

3.1.6

mission phase

period of a mission during which specified communications characteristics are fixed.

Note 1 to entry: The transition between two consecutive mission phases may cause an interruption of the communications services.

3.1.7

post-mission orbit lifetime

duration of the orbit after completion of all mission phases

Note 1 to entry: The disposal phase duration is a component of post-mission duration.

3.1.8**space object**

man-made object in outer space

3.1.9**orbit**

path followed by a space object

3.1.10**solar cycle**

≈11-year solar cycle based on the 13-month running mean for monthly sunspot number and is highly correlated with the 13-month running mean for monthly solar radio flux measurements at the 10,7 cm wavelength

Note 1 to entry: Historical records back to the earliest recorded data (1945) are shown in [Figure 2](#).

Note 2 to entry: For reference, the 25-year post-mission orbit lifetime constraint specified in ISO 24113 is overlaid onto the historical data; it can be seen that multiple solar cycles are encapsulated by this long time duration.

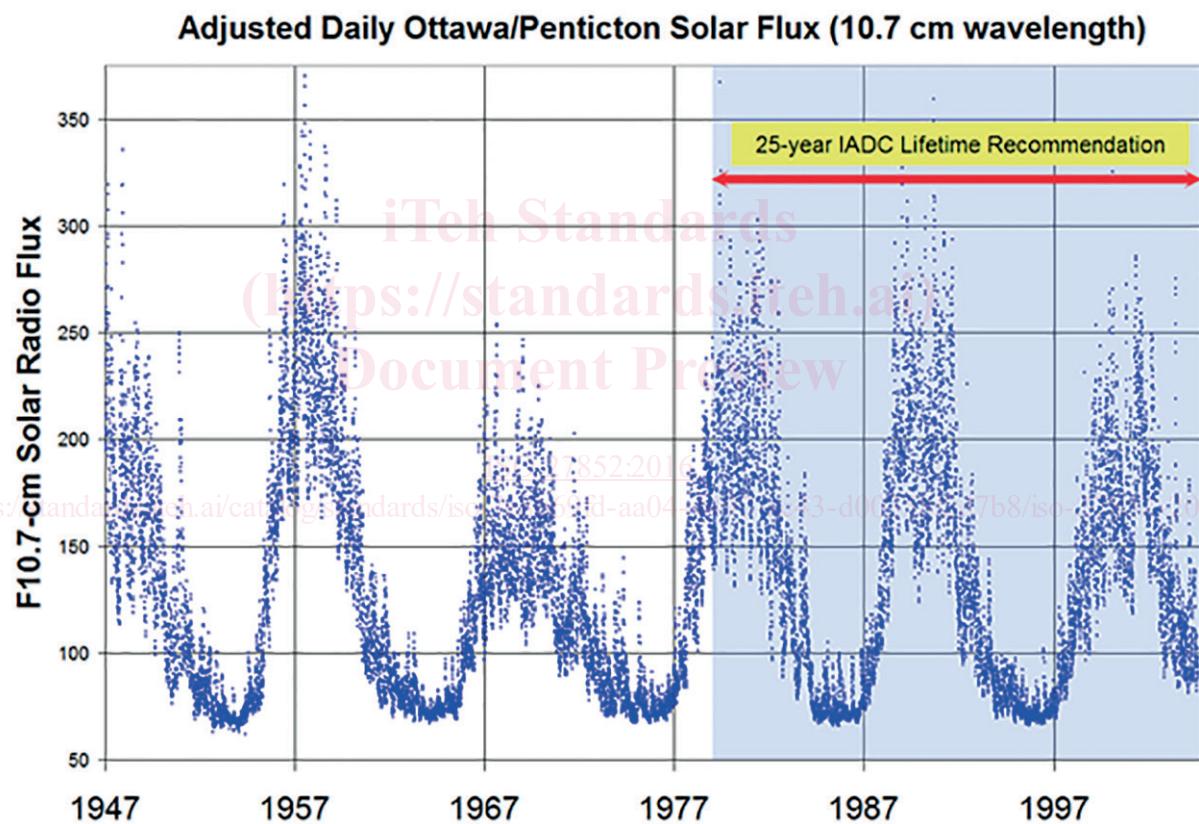


Figure 2 — Solar cycle (≈11-year duration)

3.2 Symbols

a	orbit semi-major axis
A	spacecraft cross-sectional area with respect to the relative wind
A_p	earth daily geomagnetic index
β	ballistic coefficient of spacecraft = $C_D \cdot A/m$
C_D	spacecraft drag coefficient
C_R	spacecraft reflectivity coefficient

e	orbit eccentricity
$F_{10,7}$	solar radio flux observed daily at 2 800 MHz (10,7 cm) in solar flux units ($10^{-22} \text{W m}^{-2} \text{Hz}^{-1}$)
$F_{10,7} \text{ Bar}$	solar radio flux at 2 800 MHz (10,7 cm), averaged over three solar rotations
H_a	apogee altitude = $a (1 + e) - R_e$
H_p	perigee altitude = $a (1 - e) - R_e$
m	mass of spacecraft
R_e	equatorial radius of the Earth

3.3 Abbreviated terms

3Bdy	third-body (perturbations)
CAD	computer-aided design
GEO	geosynchronous earth orbit
GTO	geosynchronous transfer orbit
HAMR	high area-to-mass ratio
IADC	Inter-Agency Space Debris Coordination Committee
ISO	International Organization for Standardization
LEO	low earth orbit
N/A	not applicable
RAAN	orbit right ascension of the ascending node (angle between vernal equinox and orbit ascending node, measured CCW in equatorial plane, looking in-Z direction)
SRP	solar radiation pressure
STSC	Scientific and Technical Subcommittee of the Committee
UNCOPUOS	United Nations Committee on the Peaceful Uses of Outer Space

4 Orbit lifetime estimation

[ISO 27852:2016](https://standards.iteh.ai/catalog/standards/iso/7e8b69fd-aa04-4497-ac43-d00358f2d7b8/iso-27852-2016)

4.1 General requirements

The orbital lifetime of LEO-crossing mission-related objects shall be estimated using the processes specified in this International Standard. In addition to any user-imposed constraints, the post-mission portion of the resulting orbit lifetime estimate shall then be constrained to a maximum of 25 years per ISO 24113 using a combination of (a) initial orbit selection, (b) spacecraft vehicle design, (c) spacecraft launch and early orbit concepts of operation which minimize LEO-crossing objects, (d) spacecraft ballistic parameter modifications at EOL, and (e) spacecraft deorbit maneuvers.

4.2 Definition of orbit lifetime estimation process

The orbit lifetime estimation process is represented generically in [Figure 3](#).

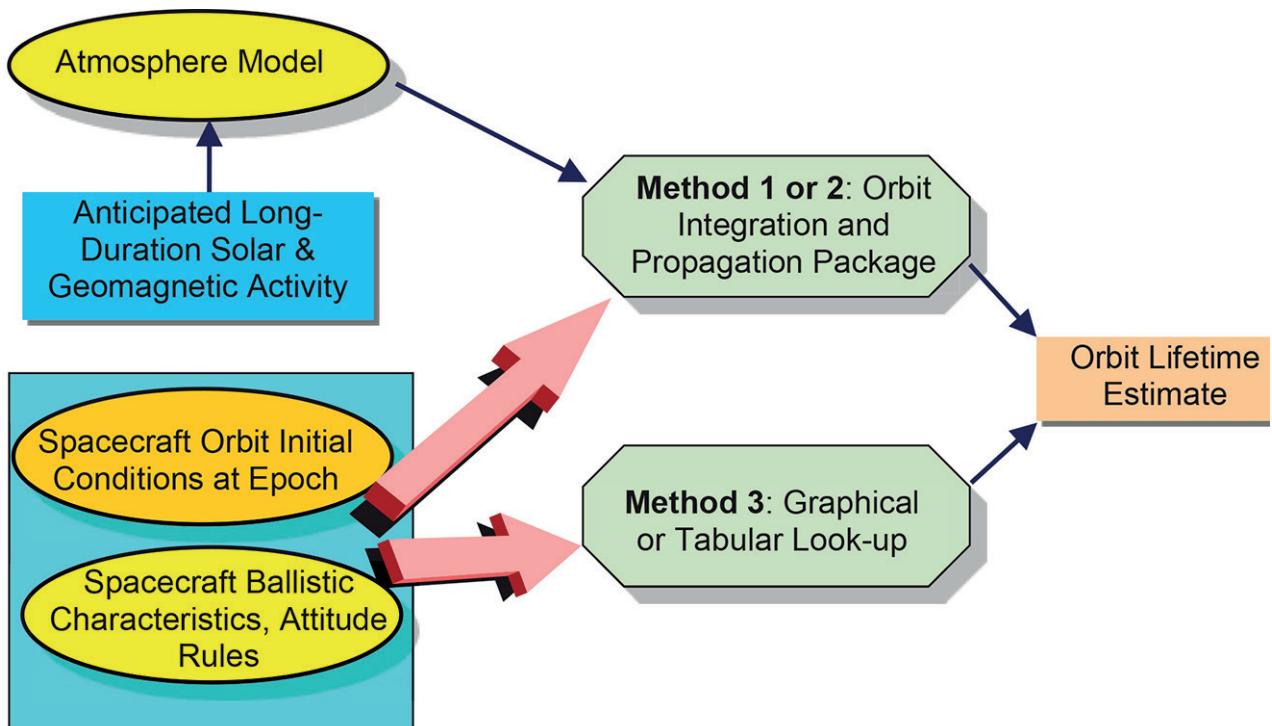


Figure 3 — Orbit lifetime estimation process^[4]

(<https://standards.iteh.ai>)

5 Orbit lifetime estimation methods and applicability

5.1 General

There are three basic analysis methods used to estimate orbit lifetime,^[3] as depicted in [Figure 3](#).
 Determination of the method used to estimate orbital lifetime for a specific space object shall be based upon the orbit type and perturbations experienced by the spacecraft as shown in [Table 1](#).

Table 1 — Applicable method with mandated conservative margins of error (in percent) and required perturbation modelling

Orbit apogee altitude, km	Special orbit		Conservative margin applied to each method			
	Sun-sync?	High area-to-mass?	Method 1: Numerical integration	Method 2: Semi-analytic	Method 3: Table look-up	Method 3: Graph, formula fit
Apogee < 2 000 km	No	No	No margin req'd	5 % margin	10 % margin	25 % margin
Apogee < 2 000 km	No	Yes	No margin; use SRP	5 % margin; use SRP	10 % margin IFF $C_r \approx 1,7$	N/A
Apogee < 2 000 km	Yes	No	No margin req'd	5 % margin	N/A	N/A
Apogee < 2 000 km	Yes	Yes	No margin req'd; use SRP	5 % margin; use SRP	N/A	N/A
Apogee > 2 000 km	Either	Either	No margin req'd; use 3Bdy+SRP	5 % margin; use 3Bdy+SRP	N/A	N/A

N/A = not applicable
3Bdy = third-body perturbations
SRP = solar radiation pressure

Method 1, certainly the highest fidelity model, utilizes a numerical integrator with a detailed gravity model, third-body effects, solar radiation pressure, and a detailed spacecraft ballistic coefficient model. Method 2 utilizes a definition of mean orbital elements,^{[4] [5]} semi-analytic orbit theory and average spacecraft ballistic coefficient to permit the very rapid integration of the equations of motion while still retaining reasonable accuracy. Method 3 is simply a table lookup, graphical analysis or evaluation of formulae that have been fit to pre-computed orbit lifetime estimation data obtained via the extensive and repetitive application of Methods 1 and/or 2. It is worth noting that all methods (1 through 3) shall include at gravity zonals J_2 and J_3 at a minimum.

5.2 Method 1: High-precision numerical integration

Method 1 is the direct numerical integration of all accelerations in Cartesian space, with the ability to incorporate a detailed gravity model (e.g. using a larger spherical harmonics model to address resonance effects), third-body effects, solar radiation pressure, vehicle attitude rules or aero-torque-driven attitude torques, and a detailed spacecraft ballistic coefficient model based on the variation of the angle-of-attack, with respect to the relative wind. Atmospheric rotation at the Earth's rotational rate is also easily incorporated in this approach. The only negative aspects to such simulations are (a) they run much slower than Method 2, (b) many of the detailed data inputs required to make this method realize its full accuracy potential are simply unavailable, and (c) any gains in orbit lifetime prediction accuracy are frequently overwhelmed by inherent inaccuracies of atmospheric modelling and associated inaccuracies of long term solar activity predictions/estimates. However, to analyse a few select cases where such detailed model inputs are known, this is undoubtedly the most accurate method. At a minimum, Method 1 orbit lifetime estimations shall account for J_2 and J_3 perturbations and drag using an accepted atmosphere model and an averaged ballistic coefficient. In the case of high apogee orbits (e.g. geosynchronous transfer orbits) or other resonant orbits, sun and moon third-body perturbations and solar radiation pressure effects shall also be modelled (see Reference [28] for additional discussion).

5.3 Method 2: Rapid semi-analytical orbit propagation

Method 2 analysis tools utilize semi-analytic propagation of mean orbit elements^[4] ^[5] influenced by gravity zonals J_2 and J_3 and selected atmosphere models. The primary advantage of this approach over direct numerical integration of the equations of motion (Method 1) is that long-duration orbit lifetime cases can be quickly analysed (e.g. 1 s versus 1 700 s CPU time for a 30-year orbit lifetime case). While incorporation of an attitude-dependent ballistic coefficient is possible for this method, an average ballistic coefficient is typically used. At a minimum, Method 2 orbit lifetime estimations shall account for J_2 and J_3 perturbations and drag using an accepted atmosphere model and an average ballistic coefficient. In the case of high apogee orbits (e.g. GTO), sun and moon third-body perturbations shall also be modelled.

5.4 Method 3: Numerical table look-up, analysis and fit formula evaluations

In this final method, one uses tables, graphs and formulae representing data that was generated by exhaustively using Methods 1 and 2 (see [5.2](#) and [5.3](#)). The graphs and formulae provided in this International Standard can help the analyst crudely estimate orbit lifetime for their particular case of interest; the electronic access to tabular look-up provided via this International Standard (at www.Celestrak.com) permits the analyst to estimate orbit lifetime for their particular case of interest via interpolation of Method 1 or Method 2 gridded data; all such Method 3 data in this International Standard were generated using Method 2 approaches. At a minimum, Method 3 orbit lifetime products shall be derived from Method 1 or Method 2 analysis products meeting the requirements stated above. When using this method, the analyst shall impose at least a 10 % margin of error to account for table look-up interpolation errors. When using graphs and formulae, the analyst shall impose a 25 % margin of error.

5.5 Orbit lifetime sensitivity to sun-synchronous

For sun-synchronous orbits, orbit lifetime has some sensitivity to the initial value of RAAN due to the density variations with the local sun angle. Results from numerous orbit lifetime estimations show that orbits with 6:00 am local time have longer lifetime than orbits with 12:00 noon local time by about 5,5 %. ^[3] This maximum difference (500 d) translates into a 5 % error which can be corrected by knowing the local time of the orbit. As a result, Methods 1 or 2 analyses of the actual sun-synchronous orbit condition shall be used when estimating the lifetime of sun-synchronous orbits (see References [\[28\]](#) and [\[38\]](#), where more details are given).

5.6 Orbit lifetime statistical approach for high-eccentricity orbits (e.g. GTO)

For high-eccentricity orbits (particularly geosynchronous transfer orbits or GTO), it can be difficult to iterate to lifetime threshold constraints due to the coupling in eccentricity between the third-body perturbations and the drag decay. Due to this convergence difficulty, only Method 1 or 2 analyses shall be used when determining initial conditions which achieve a specified lifetime threshold for such orbits.

Sample analyses of GTO launcher stages (see References [\[29\]](#) and [\[30\]](#)) highlight this orbit lifetime sensitivity to initial conditions (orbit, spacecraft characteristic and force model), leading to a wide spectrum of orbital lifetimes.

Some theoretical considerations about the dynamical properties of GTO orbits are provided in References [\[29\]](#) and [\[36\]](#).

The following test case illustrates the complex dynamical properties of GTO. Initial parameters are provided in [Table 2](#).

Table 2 — GTO initial conditions for the Monte Carlo simulation

Perigee altitude	200 km
Apogee altitude	GEO altitude
Inclination	2°

Table 2 (continued)

Area to mass ratio	5e-3 m ² /kg
Solar activity	Constant (F10.7 = 140 sfu Ap = 15)
Drag coefficient	Constant = 2,2
Reflectivity coefficient	Constant = 2

[Figure 4](#) shows lifetime results (years) when varying the initial date and the initial local time of perigee. This latest parameter is defined as the angle in the equator between the sun direction and the orbit perigee, measured in hours. The date was chosen from day 1 to 365 in year 1998 and the local time of perigee was chosen by varying the right ascension of ascending node from 0π to 2π . A total of 2 500 different initial conditions were generated.

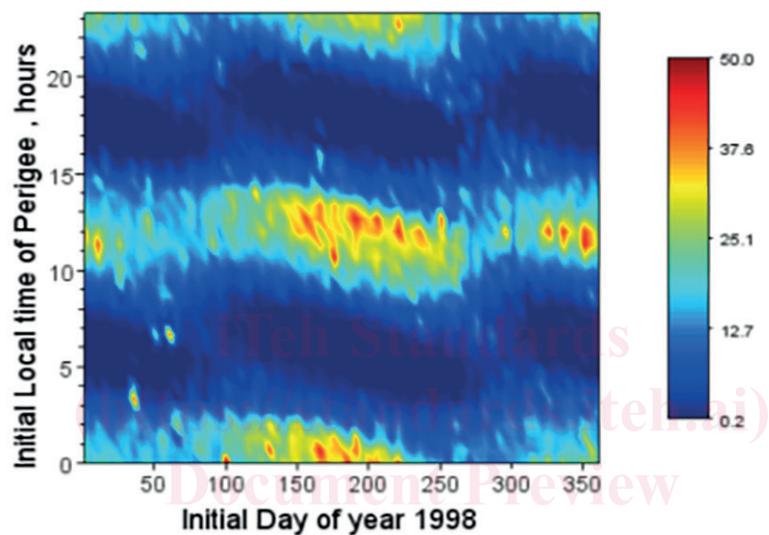


Figure 4 — Lifetime variations with respect to initial date and local time of perigee (year)

The shapes of the lifetime contours confirm that initial day of year and local time of perigee are initial conditions that make sense to describe GTO evolution since strong patterns are visible. The amplitudes of lifetimes variations are worth noting: from several months to more than 50 years. Previous results (see References [30] and [37]) are illustrated here: the longest lifetimes are obtained for initial sun-pointing (12 h local time) or anti sun-pointing (24 h local time) perigee with an initial date around the solstices. Note that the dark red pixels drawn in dark blue areas, as seen for initial day 60 and local time 7 h, are an indication of the presence of strong resonance phenomena. We know that the year also has an influence, to a lesser extent, through the moon perturbation.

[Figure 5](#) shows semi-major axis evolution for several propagations of a typical low-inclined GTO. The different curves correspond to changes of 0,1 % or 1 % in the area to mass ratio of the object (A/m), which is far below the level of incertitude on this parameter. These dispersions lead to variations of decades in the re-entry duration. Such a strong non-linear behaviour is explained by the aforementioned resonances. One can see that semi-major axis evolutions are quite similar between all propagation cases until the entrance in the coupling between J2 and sun perturbations, for a semi-major axis equal to about 15 500 km. The duration of the resonance (period when the semi-major axis remains constant) and, thus, the rest of the propagation are completely different. A similar figure can be plotted by keeping the area to mass ratio constant and slightly changing the solar activity.